

From:Landau Fermi liquid & two-fluid hydrodynamicsTo:physics of quantum vacuum & cosmology

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1. effective (phenomenological) theories of hydrodynamic type

Landau two-fluid hydrodynamics

Einstein general relativity



2. effective (phenomenological) theories from p-space topology



Landau theory of Fermi liquid

Standard Model + gravity





quantum vacuum as Lorentz invariant medium application to cosmology



3+1 sources of effective (**phenomenological**) theory of quantum liquids & relativistic QFT



1. Effective theories of hydrodynamic type

Landau two-fluid hydrodynamics

Einstein general relativity

classical low-energy property of quantum liquids

Landau equations



classical low-energy property of quantum vacuum

Einstein equations



Landau equations & Einstein equations are effective theories describing dynamics of metric field + matter (quasiparticles)



Landau quasiparticles

weakly excited state can be considered as system of "elementary excitations" Landau, 1941



N.G. Berloff & P.H. Roberts Nonlocal condensate models of superfluid helium J. Phys. A32, 5611 (1999)



long-wave quasiparticles: `relativistic' phonons

effective metric

$$E(k) = ck$$

c - speed of sound
or light

$$g^{\mu\nu}k_{\mu}k_{\nu} = 0$$

long-wave quasiparticles: relativistic fermions & bosons



general relativity

Effective metric in Landau two-fluid hydrodynamics



Landau critical velocity = black hole horizon



Superfluids	_	Universe
acoustic gravity met	Theories of gravi	ty general relativity
geometry of effective space tin for quasiparticles (phonons) geodesics for phonons Landau two-fluid equations	$ds^{2} = g_{\mu\nu}dx^{\mu}dx^{\nu} = 0$	geometry of space time for matter geodesics for photons Einstein equations of GR
$\dot{\mathbf{\rho}} + \nabla \cdot (\rho \mathbf{v}_{s} + \mathbf{P}^{Matter}) = 0$ $\dot{\mathbf{v}}_{s} + \nabla (\mu + \mathbf{v}_{s}^{2}/2) = 0$	dynamic equations for metric field $g_{\mu\nu}$	$\frac{1}{8\pi G} \left(R_{\mu\nu} - g_{\mu\nu} R/2 \right) = T_{\mu\nu}^{Matter}$
equationsequationfor superfluidfor normalcomponentcomponent	$T^{\mu\nu}_{;\nu Matter} = 0$ for	quation or matter 1/2 of GR



2. Effective theories from momentum space topology



Standard Model + gravity

two major universality classes of fermionic vacua: vacuum with Fermi surface & vacuum with Fermi point



crossover from Landau 2-fluid hydrodynamics to Einstein general relativity they represent two different limits of hydrodynamic type equations

> equations for $g^{\mu\nu}$ depend on hierarchy of ultraviolet cut-off's: Planck energy scale E_{Planck} vs Lorentz violating scale E_{Lorentz}



$$\begin{split} & E_{\text{Planck}} >> E_{\text{Lorentz}} \\ & \text{Landau two-fluid} \\ & \text{hydrodynamic equations} \end{split}$$

 $E_{\text{Planck}} \ll E_{\text{Lorentz}}$

Einstein equations of general relativity



³He-A with Fermi point

Universe

 $E_{\rm Lorentz} << E_{\rm Planck}$ $E_{\rm Lorentz} \sim 10^{-3} E_{\rm Planck}$

 $E_{\text{Lorentz}} \gg E_{\text{Planck}}$ $E_{\text{Lorentz}} > 10^{15} E_{\text{Planck}}$



high-energy physics and cosmology are extremely ultra-low temperature physics

characteristic high-energy scale in our vacuum (analog of atomic scale in cond-mat) is Planck energy

 $E_{\rm P} = (hc^5/G)^{1/2} \sim 10^{19} \,{\rm GeV} \sim 10^{32} {\rm K}$

highest energy in accelerators $E_{\rm ew} \sim 1 {
m TeV} \sim 10^{16} {
m K}$

$$E_{\rm ew} \sim 10^{-16} E_{\rm Planck}$$



cosmology is extremely ultra-low frequency physics

v(r) = Hr

B. L. Hu New View on Quantum Gravity and the Origin of the Universe

gr-qc/0611058



Why no freezing at low T?

natural masses of elementary particles are of order of characteristic energy scale the Planck energy

$$m \sim E_{\text{Planck}} \sim 10^{19} \text{ GeV} \sim 10^{32} \text{K}$$

even at highest temperature we can reach

$T \sim 1 \text{ TeV} \sim 10^{16} \text{K}$

everything should be completely frozen out











reason:

quasiparticles leaving near Fermi surface have no gap

quasiparticles leaving near Fermi point have no mass

equilibrium condition in theories of hydrodynamic type: Landau 2-fluid hydrodynamics & hydrodynamics of vacuum Why no freezing at low T?









Universality classes of quantum vacua

physics at low T is determined by low-lying excitations





topology *protects vortices* & *hedgehogs:* **one cannot comb the hair on a ball smooth**





Route to Landau Fermi-liquid





From Landau Fermi-liquid to Standard Model From Fermi surface to Fermi point



 p_{z} p_{z} $p_{\rm v}$ p_x p_x hedgehog with spines (spins) hedgehog with spines (spins) inward $(N_3 = -1)$ **outward** ($N_3 = +1$) $H = +c \sigma \cdot \mathbf{p}$ $H = -c \, \boldsymbol{\sigma} \cdot \mathbf{p}$ right $H = \boldsymbol{\sigma} \cdot \mathbf{g}(\mathbf{p})$ left $\mathbf{g}(\mathbf{p}) = -\mathbf{c}\mathbf{p}$ neutrino $\mathbf{g}(\mathbf{p}) = +\mathbf{c}\mathbf{p}$ neutrino

Topological invariant for right-handed and left-handed elementary particles

$$N_{3} = \frac{1}{8\pi} e_{ijk} \int dS^{i} \mathbf{\hat{g}} \cdot (\mathbf{\partial}^{j} \mathbf{\hat{g}} \times \mathbf{\partial}^{k} \mathbf{\hat{g}})$$

over 2D surface
around Fermi point





Chiral fermions in Standard Model

Family #1 of quarks and leptons



examples of Fermi points in condensed matter

superfluids & superconductors with point nodes in gap: superfluid ³He-A, chiral superconductor Sr₂RuO₄, triplet cold Fermi gases Gap node - Fermi point (anti-hedgehog) $N_3 = -1$ E $N_3 = \frac{1}{8\pi} e_{ijk} \int d\mathbf{S}^k \, \hat{\mathbf{g}} \cdot (\partial_{p_i} \, \hat{\mathbf{g}} \times \partial_{p_j} \, \hat{\mathbf{g}})$ over 2D surface S in 3D p-space $N_3 = 1$ Gap node - Fermi point (hedgehog) 16

emergence of relativistic particles

original non-relativistic Hamiltonian

$$H = \begin{pmatrix} \frac{p^2}{2m} - \mu & c(p_x + ip_y) \\ c(p_x - ip_y) & -\frac{p^2}{2m} + \mu \end{pmatrix} = \begin{pmatrix} g_3(\mathbf{p}) & g_1(\mathbf{p}) + i g_2(\mathbf{p}) \\ g_1(\mathbf{p}) - i g_2(\mathbf{p}) & -g_3(\mathbf{p}) \end{pmatrix} = \mathbf{\tau} \cdot \mathbf{g}(\mathbf{p})$$
close to nodes, i.e. in low-energy corner relativistic chiral fermions emerge
$$H = N_3 c \, \mathbf{\tau} \cdot \mathbf{p}$$

$$E = \pm cp$$

$$k = \pm cp$$

$$k = transformed equation (transformed equation ($$

bosonic collective modes in two generic fermionic vacua



two generic quantum field theories of interacting bosonic & fermionic fields

relativistic quantum fields and gravity emerging near Fermi point

Atiyah-Bott-Shapiro construction:

linear expansion of Hamiltonian near the nodes in terms of Dirac Γ -matrices



Extension of Landau idea: electric charge purely from vacuum polarization

Landau proposal for QED

Landau, 1955

$$\frac{1}{e^2} = \frac{v}{3\pi} \ln \left(\frac{E_{\text{Planck}}^2}{m^2} \right)$$

v number of charged particles E_{Planck} Planck energy *m* electron mass

v = 12

$$N_F = 4$$

extension to Standard Model which as effective theory has two different cut-off for bosons & fermions



$$\frac{1}{e^2} = \frac{8N_F}{9\pi} \ln \left(\frac{E_{\rm UV}^2}{m_Z^2} \right) - \frac{11}{6\pi} \ln \left(\frac{E_{\rm Planck}^2}{m_Z^2} \right)$$

 N_F number of families $E_{\rm UV}$ ultraviolet cutoff Klinkhamer-Volovik JETP Lett. **81** (2005) 551

 $m_{\rm Z}$ mass of Z-boson

physics at the intermediate Planck scale is Lorentz invariant





crossover from Landau 2-fluid hydrodynamics to Einstein general relativity they represent two different limits of hydrodynamic type equations

> equations for $g^{\mu\nu}$ depend on hierarchy of ultraviolet cut-off's: Planck energy scale E_{Planck} vs Lorentz violating scale E_{Lorentz}



 $E_{\text{Planck}} >> E_{\text{Lorentz}}$ emergent Landau
two-fluid hydrodynamics

³He-A with Fermi point

 $E_{\text{Planck}} << E_{\text{Lorentz}}$ emergent general covariance & general relativity



Universe

 $E_{\text{Lorentz}} \ll E_{\text{Planck}} \qquad E_{\text{Lorentz}} \gg E_{\text{Planck}}$ $E_{\text{Lorentz}} \sim 10^{-3} E_{\text{Planck}} \qquad E_{\text{Lorentz}} > 10^{9} E_{\text{Planck}}$

3. Extension of Landau ideas

vacuum as Lorentz invariant medium & application to cosmology



Cosmological constant - weight of ether (vacuum vacuum)

"ether should not be thought of as endowed with the quality characteristic of ponderable media ... The idea of motion may not be applied to it."

wait !

Einstein, 1920

zero cosmological constant ???



perfect vacuum is weightless

"ether should not be thought of as endowed with the quality characteristic of ponderable media ... The idea of motion may not be applied to it."

Einstein, 1920



Lorentz invariance is guiding principle





extension of Landau ideas: vacuum as Lorentz invariant medium & application to cosmology

vacuum`empty space' $P_{vac} = -\tilde{\varepsilon}_{vac}$ \blacktriangleleft external pressure P

 $\Lambda = \tilde{\varepsilon}_{vac} = -P_{vac}$ energy density pressure of vacuum of vacuum

 $P_{\rm vac} = - dE/dV = - \tilde{\varepsilon}_{\rm vac}$

$$\chi_{\rm vac} = -(1/V) \ dV/dP$$

compressibility of vacuum

conclusions

F.R. Klinkhamer, G.E. Volovik Self-tuning vacuum variable & cosmological constant PRD 77, 085015 (2008); Dynamic vacuum variable & equilibrium approach in cosmology arxiv: 0806.2805

$$<(\Delta P_{\rm vac})^2 > = T/(V\chi_{\rm vac})$$
$$<(\Delta \Lambda)^2 > = <(\Lambda P)^2 >$$

pressure fluctuations

natural value of Λ determined by macroscopic physics

 $\Lambda \sim 0$

natural value of χ_{vac} determined by microscopic physics

$$\chi_{\rm vac} \sim E_{\rm Planck}^{-4}$$

volume of Universe is large:

 $V > T_{\rm CMB} / (\Lambda^2 \chi_{\rm vac})$





thermodynamics & dynamics of Lorentz invariant vacuum

energy density $\epsilon_{vac} \left(u^{\mu}_{v} \right)$ of vacuum is function of

equilibrium vacuum is obtained from equation

$$\delta \varepsilon_{\rm vac} / \delta u^{\mu} = \nabla_{\nu} (\delta \varepsilon_{\rm vac} / \delta u^{\mu}{}_{\nu}) = 0$$

equilibrium solution:

$$u_{\mu\nu} = u g_{\mu\nu}$$
 $u = const$

$$\boldsymbol{u}^{\mu}_{\nu} = \boldsymbol{\nabla}_{\!\!\nu} \boldsymbol{u}^{\mu}$$

microscopic vacuum energy

$$\Lambda = \varepsilon_{\text{vac}}(u) \sim E_{\text{Planck}}^4???$$

naive estimation of vacuum energy & cosmological constant highly disagrees with observations (Cosmological Constant Problem)

macroscopic vacuum energy: from energy momentum tensor

$$T_{\mu\nu} = \delta S / \delta g^{\mu\nu} = (\varepsilon_{\rm vac} (u) - u \, d\varepsilon_{\rm vac} / du) g_{\mu\nu}$$

It is $T_{\mu\nu}$ which is gravitating, thus cosmological constant is $\Lambda = \varepsilon_{vac}(u) - u \, d\varepsilon_{vac}/du$ dynamics of cosmological constant: from Planck scale to present value



microscopic vacuum energy has natural Planck scale:

$$\varepsilon_{\rm vac}\left(u\right) \sim E_{\rm Planck}^{4}$$

macroscopic vacuum energy

$$\varepsilon_{\rm vac}(u) - u \, \mathrm{d}\varepsilon_{\rm vac}/\mathrm{d}u = -P_{\rm vac}$$

$$\Lambda = \varepsilon_{\rm vac} \left(u \right) - u \, \mathrm{d}\varepsilon_{\rm vac} / \mathrm{d}u = -P_{\rm vac} = 0$$

two huge quantities naturally cancel each other

due to **thermodynamics**

macroscopic energy & cosmological constant have natural zero value

$$\varepsilon(\rho) - \rho d\varepsilon/d\rho = -P$$
 is macroscopic

in the absence of environment:

$$\varepsilon(\rho) - \rho d\varepsilon/d\rho = -P = 0$$

two microscopic quantities cancel each other

due to **thermodynamics**

missing topology

Landau-Lifshitz vortex sheet suggested for rotating superfluid 4He DAN 100 (1955) 669

Landau-Lifshitz vortex sheet observed in rotating superfluid 3He-A PRL 72 (1994) 3839



topologically stable



topologically unstable towards vortex lattice



topologically stable

Helsinki NMR experiments: satellite peaks from spin waves localized in & between Landau-Lifshitz vortex sheets in ³He-A Landau-Lifshitz formula: distance b between the sheets as function of rotation velocity Ω



Conclusion

Landau two-fluid hydrodynamics & Einstein general relativity are effective hydrodynamic theories: they are two different extreme limits of parameters in underlying microscopic theory

> Landau theory of Fermi liquid & Standard Model of electroweak & strong interactions are effective theories for two major classes of fermionic vacua: vacua with Fermi surface (normal 3He and metals) & vacua with Fermi point (relativistic quantum vacuum & superfluid 3He-A)

Landau ideas first applied to quantum liquids are applicable to quantum vacuum -

the modern aether



